

IceCube Neutrino Observatory

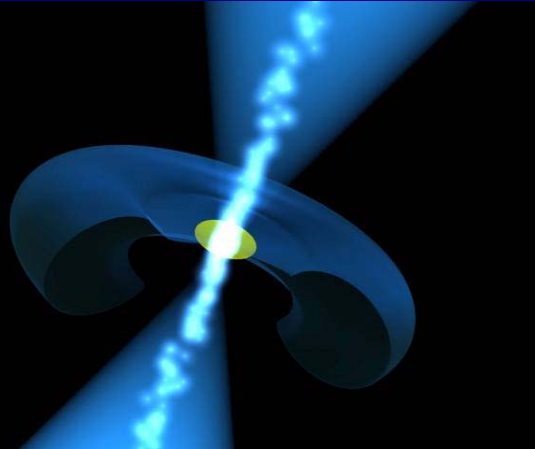
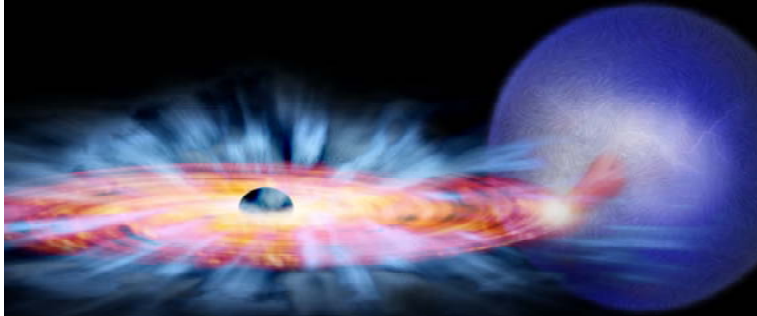


Jim Yeck
IceCube Project Director
University of Wisconsin

Why neutrino astronomy?

Astrophysical Accelerators

Accretion disk with jets



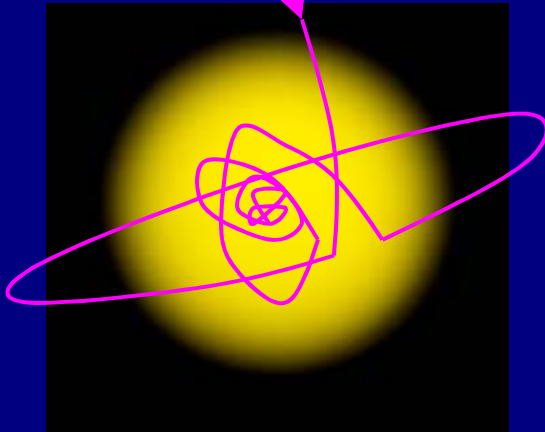
Neutrinos allow for observation of 'hidden regions' of cosmic accelerators of high energy particles.

Black holes in active galaxies, pulsars,

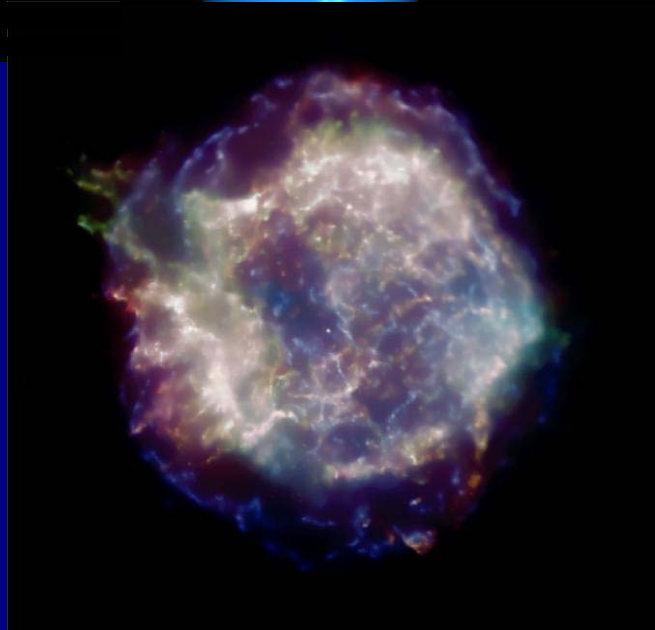
Supernova explosions,

gamma ray bursts

χ



DM annihilation

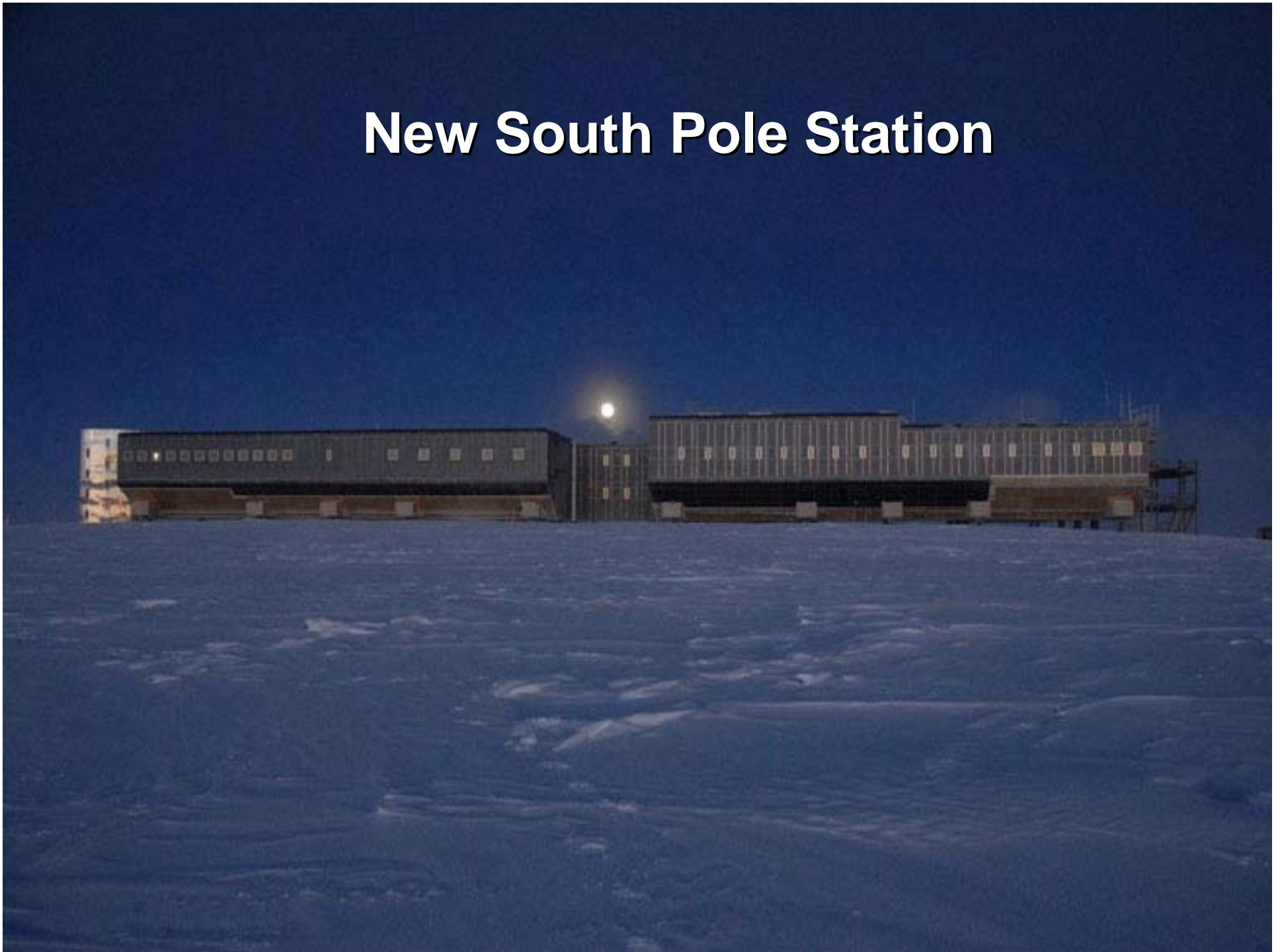


CasA Supernova Remnant in X-rays

The dome, the old station



New South Pole Station



Collaboration

<http://icecube.wisc.edu>

United states

- Univ Alaska, Anchorage
- UC Berkeley
- UC Irvine
- Clark-Atlanta University
- U Delaware / Bartol Research Inst
- University of Kansas
- Lawrence Berkeley National Lab
- University of Maryland
- Pennsylvania State University
- University of Wisconsin-Madison
- University of Wisconsin-River Falls
- Southern University, Baton Rouge



Europe

University Utrecht

- Uppsala University
- Stockholm University
- University of Oxford

- Universite Libre de Bruxelles
- Vrije Universiteit Brussel
- Université de Mons-Hainaut
- Universiteit Gent

Universität Mainz

- Humboldt Univ., Berlin
- DESY, Zeuthen
- Universität Dortmund
- Universität Wuppertal
- MPI Heidelberg
- RWTH Aachen

Japan

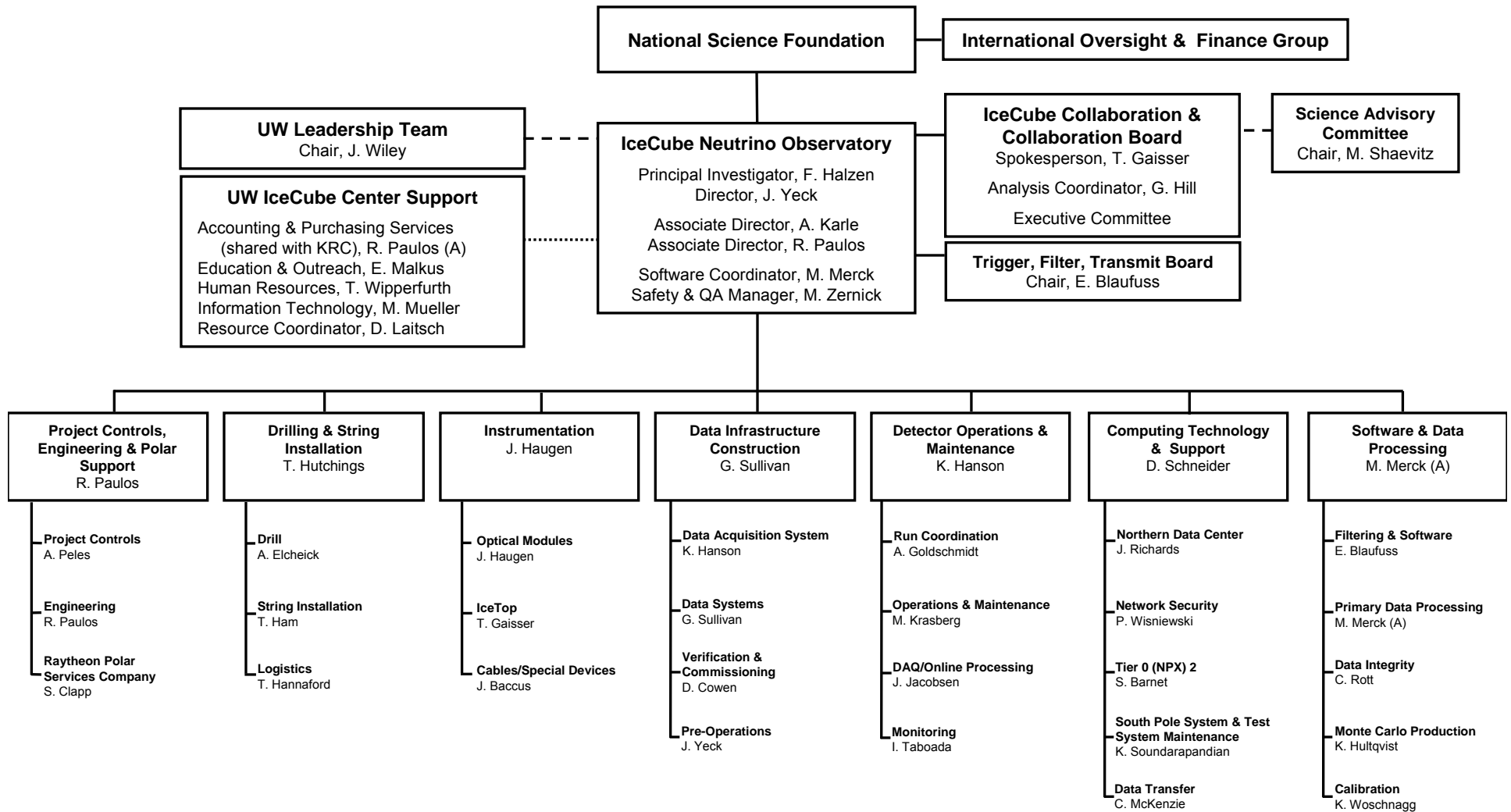
Chiba University

New Zealand

Univ. of Canterbury, Christchurch



IceCube Construction and Operations Organization



| Legend | |
|-----------|----------------------------|
| — | Direction and Reporting |
| - - - - - | Advice and Recommendations |
| | Project Support |
| (A) | Acting |

Jane H. Yeck

April 1, 2008

IceCube Detector

Deep Ice Detector

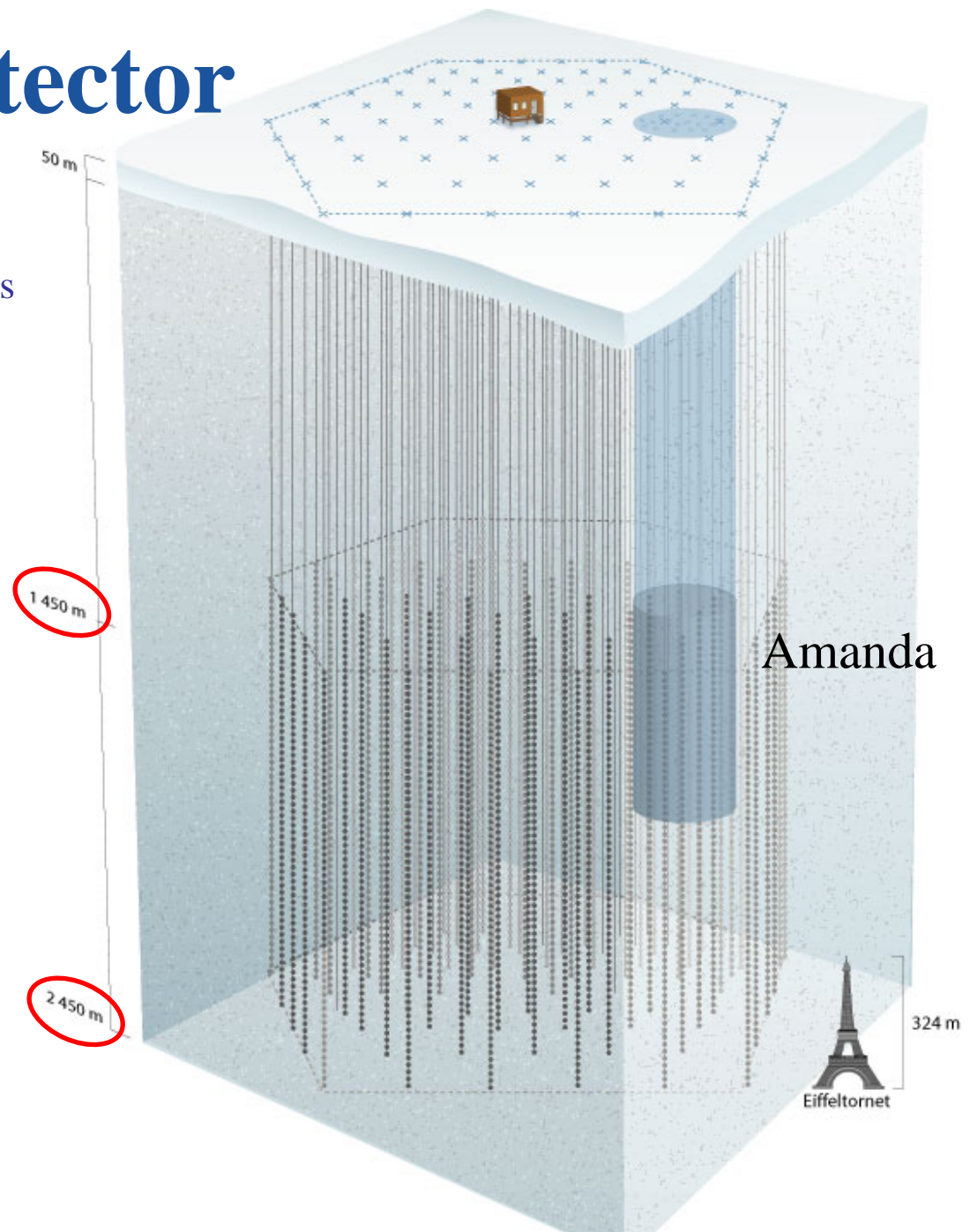
80 strings w/ 60 Digital Optical Modules
4800 Total Deep Ice DOMs
17 meters between DOMs
125 meters between strings
1 Giga Ton Detector
No single point failure in a string.

Dense “Low Energy” Core

19 String AMANDA Detector
677 Modules

IceTop Air Shower Array

80 Pairs of Cherenkov Ice Tanks
320 Total Air Shower DOMs
2 DOMs in each tank
10 meters between tanks



IceTop

80 Tanks Installed

2007-08: 18

2006-07: 13 Strings

total of 40 Strings

2005-06: 8 Strings

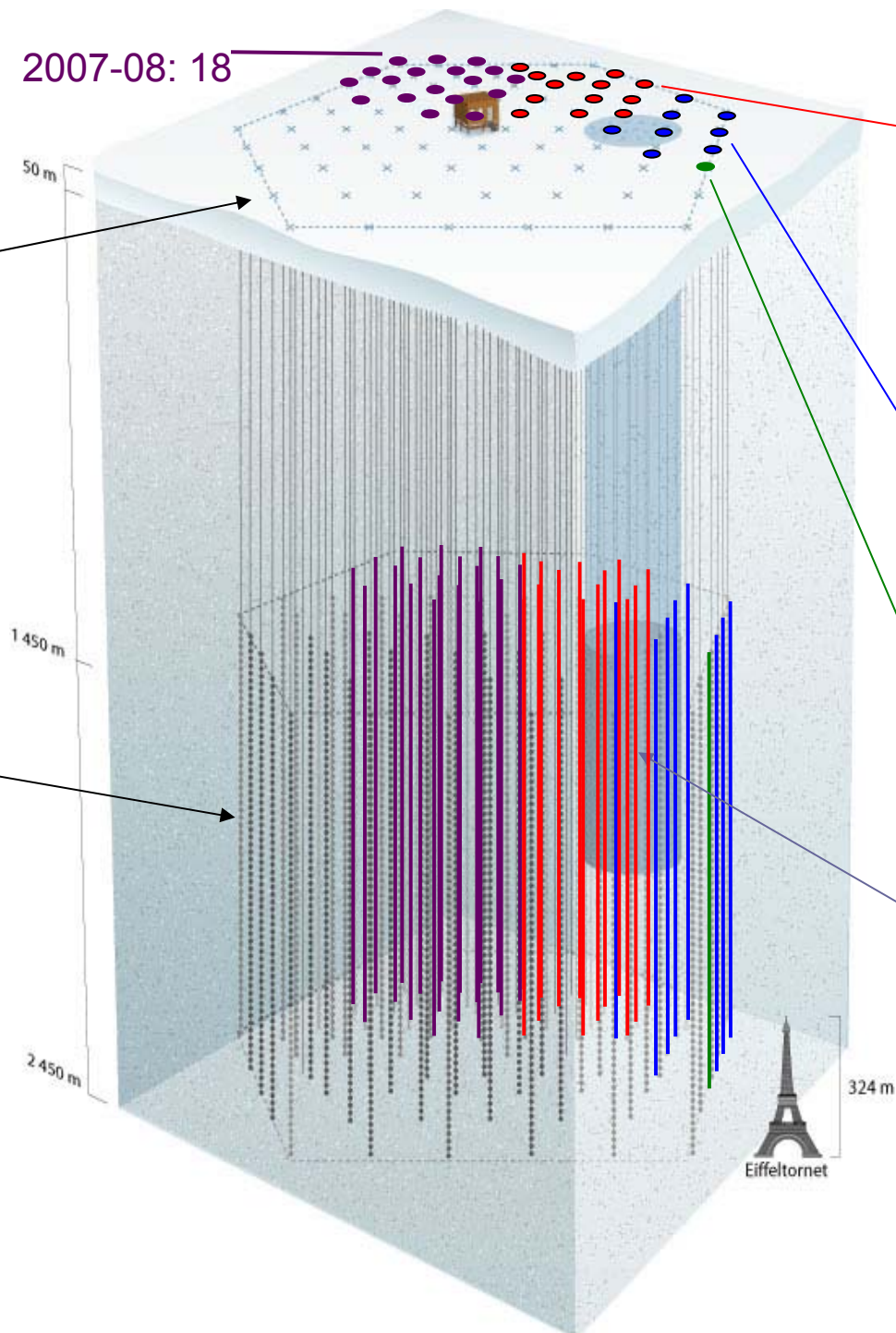
2004-05 : 1 String

*first data 2005
upgoing muon 18.
July 2005*

AMANDA
19 Strings
677 Modules

InIce

40 Strings Installed

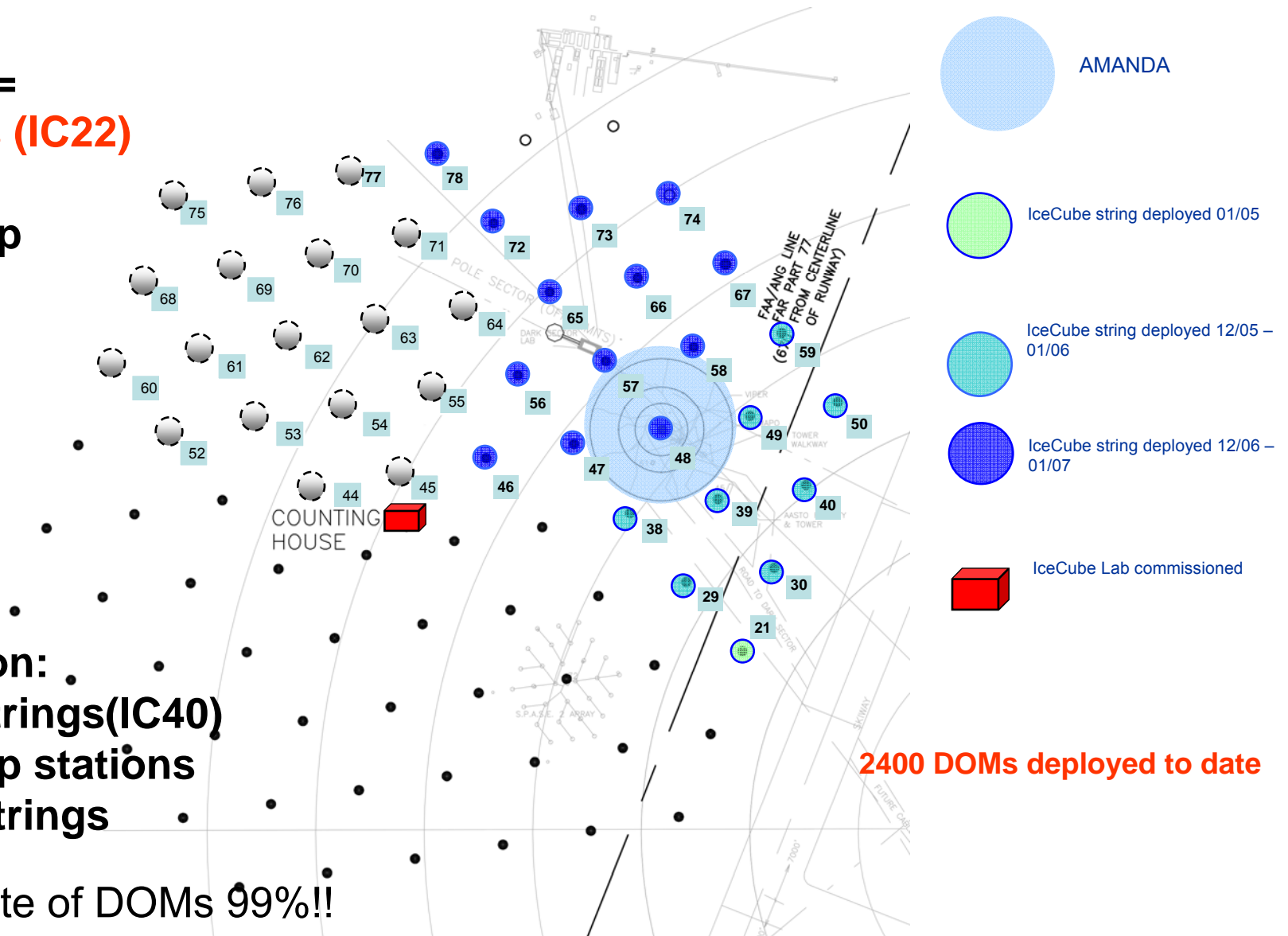


2005-2008 configurations

1 + 9 + 13 =
22 strings (IC22)
to date
+26 IceTop
stations

This season:
18 more strings (IC40)
+ 40 IceTop stations
2011: 80 strings

Survival rate of DOMs 99%!!





Original Detector Baseline

- 70 vertical strings plus surface stations
- 4,480 Digital Optical Modules (DOM)
 - 60 DOM/string spaced between 1450 and 2450 meters
 - 2 DOM/surface tank and 2 tanks at each surface location
- Instrumented volume of $\sim 1 \text{ km}^3$ of ice
- Software and computing required to commission and operate the detector

Cost & Schedule Baseline

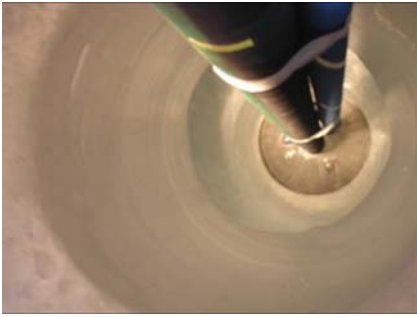
| | <u>Baseline</u> (Hartill 02/04) | <u>Current</u> (02/08) |
|------------------|---------------------------------|----------------------------|
| <u>Cost:</u> TPC | \$271.8 million | \$276.6 million |
| NSF | \$242.1 “ | \$242.1 “ <i>unchanged</i> |
| Foreign | \$ 29.7 “ | \$ 34.5 “ |

Earned Value: - \$215.5 “ (82%)

Contingency: \$40 million (22%) \$15 million (26%)

Most of the Technical Risk Retired

Schedule: 4th Quarter (Q4), 2010 Q2, 2011



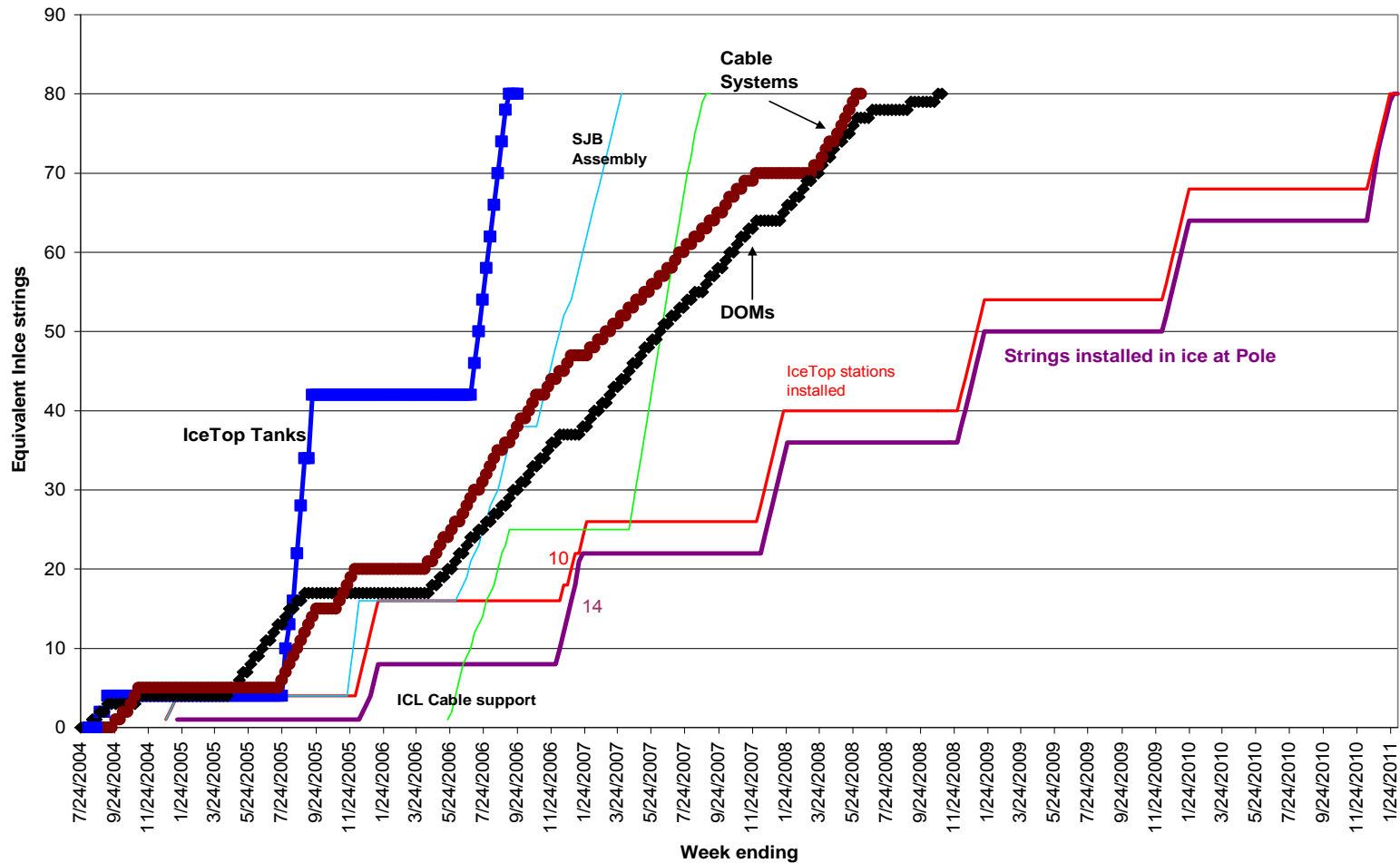
Construction Strategy

- Construction Schedule Constraints
 - South Pole construction season limited to the austral summer
 - Limits on the number of cargo and fuel flights to the South Pole
 - Limits on the number of people that can be supported at South Pole
- Critical schedule activity is the safe drilling of holes in the ice
- Maximize the instrumentation that can be installed each summer
 - Ensure that installation is not limited by instrumentation
- Provide for concurrent detector construction and operations

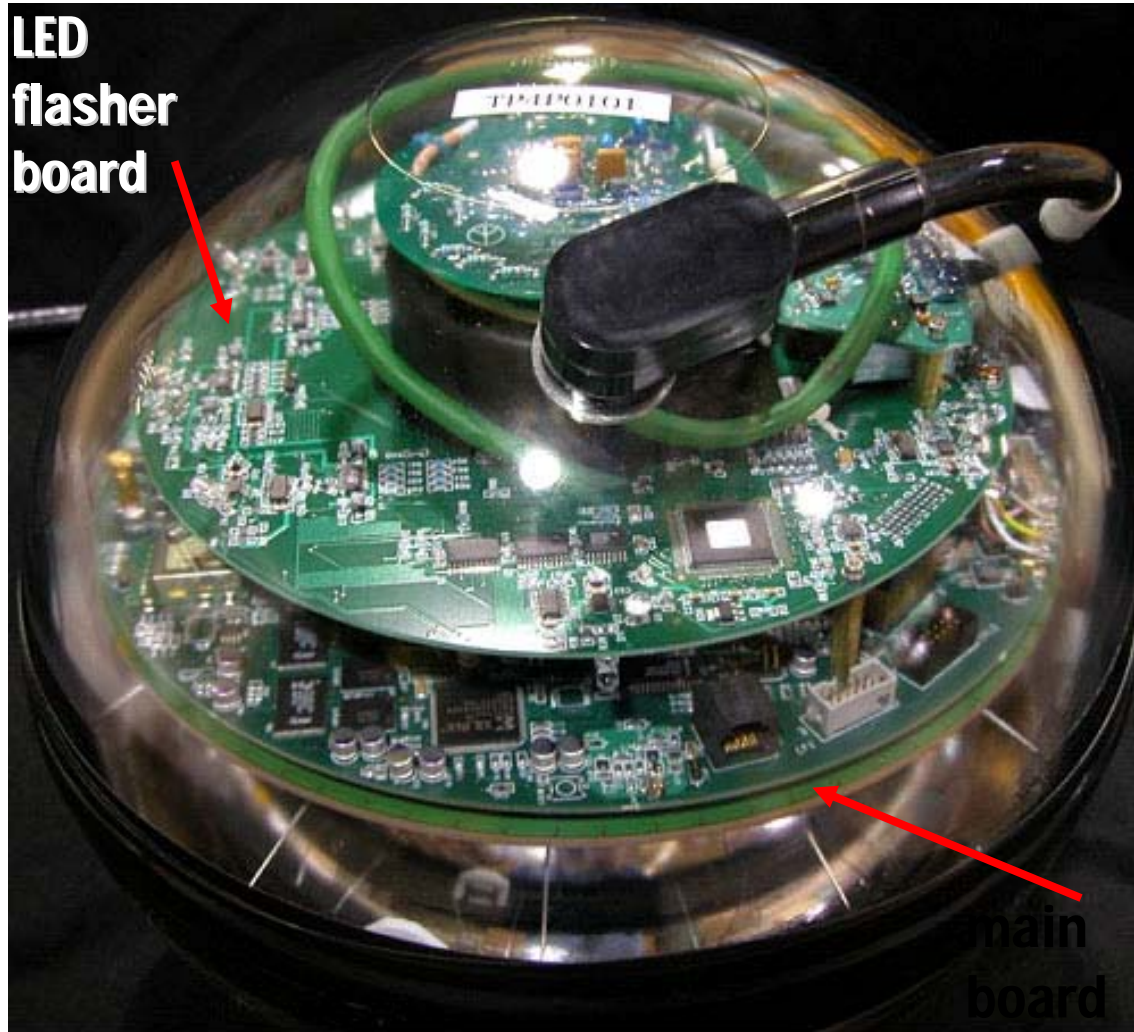


Instrumentation Production

Instrumentation Production CY2004 - CY2008 for 80 strings installed



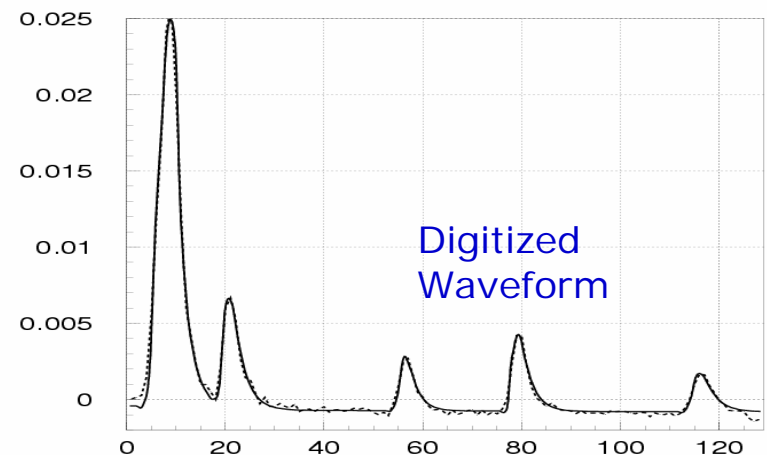
Digital Optical Module (DOM)



PMT: 10 inch Hamamatsu
Power consumption: 3 W
Digitize at 300 MHz for 400 ns with
custom chip
40 MHz for 6.4 μ s with fast ADC
Dynamic range 200pe/15 nsec

Send all data to surface over copper
2 sensors/twisted pair.
Flasherboard with 12 LEDs
Local HV

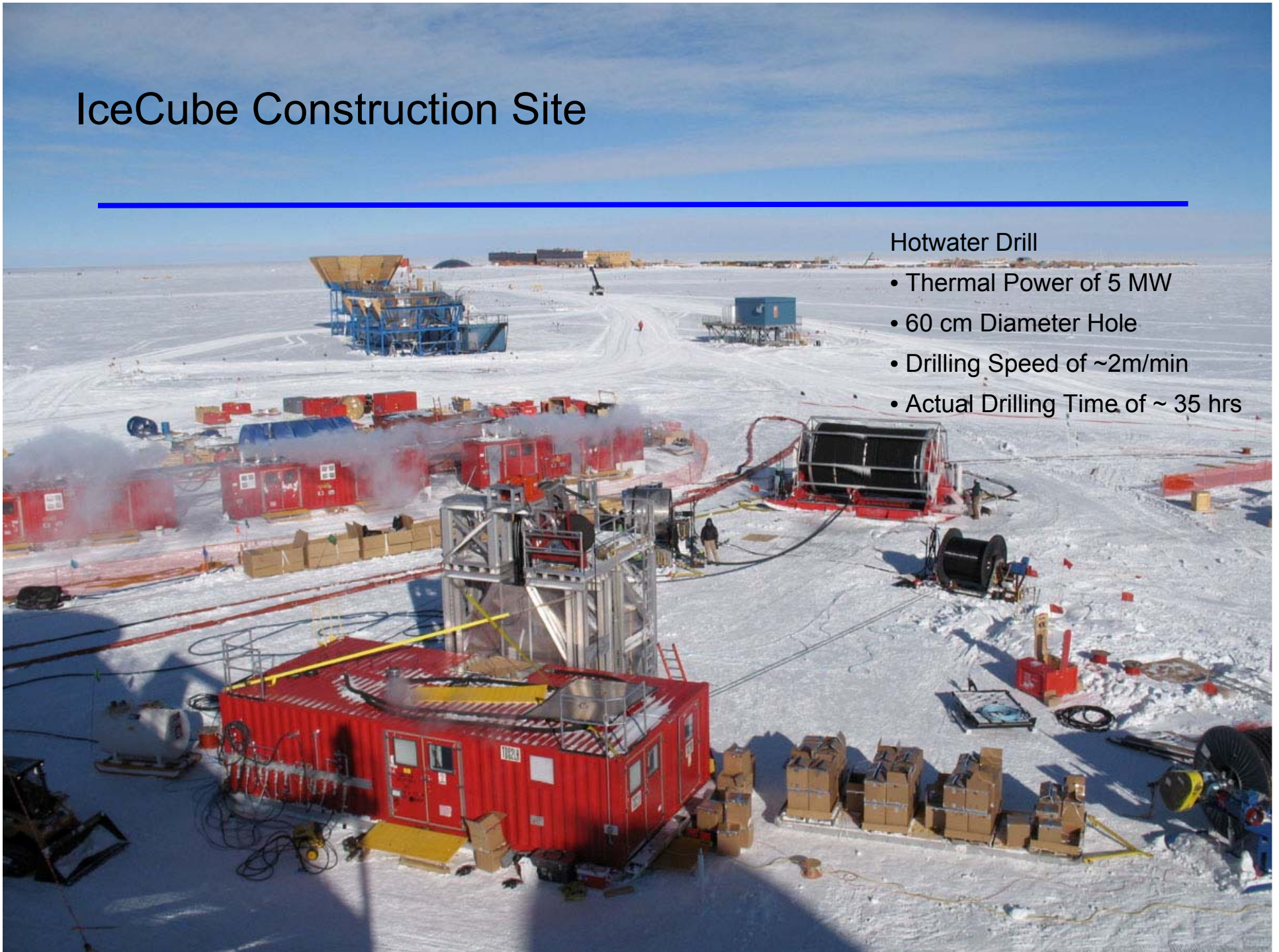
Clock stability: $10^{-10} \approx 0.1$ nsec / sec
Synchronized to GPS time every ≈ 10 sec
Time calibration resolution = 2 nsec



IceCube Construction Site

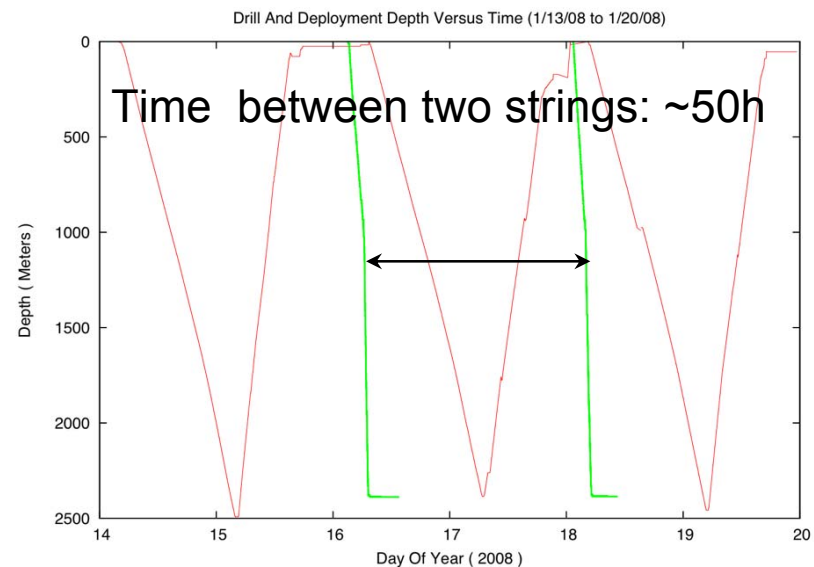
Hotwater Drill

- Thermal Power of 5 MW
- 60 cm Diameter Hole
- Drilling Speed of $\sim 2\text{m/min}$
- Actual Drilling Time of ~ 35 hrs

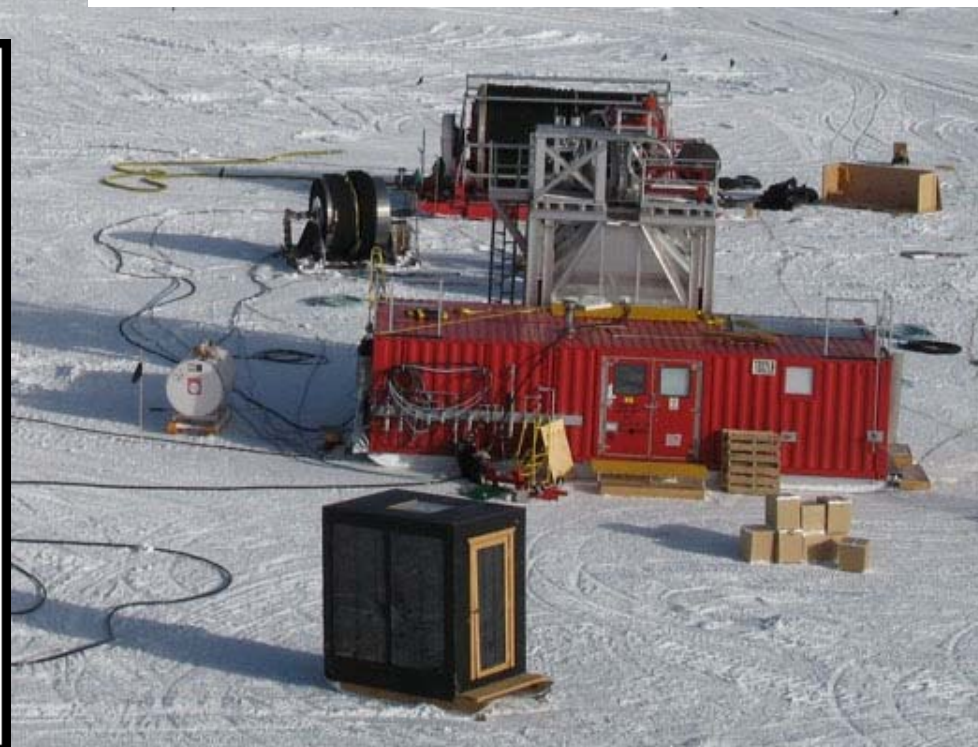
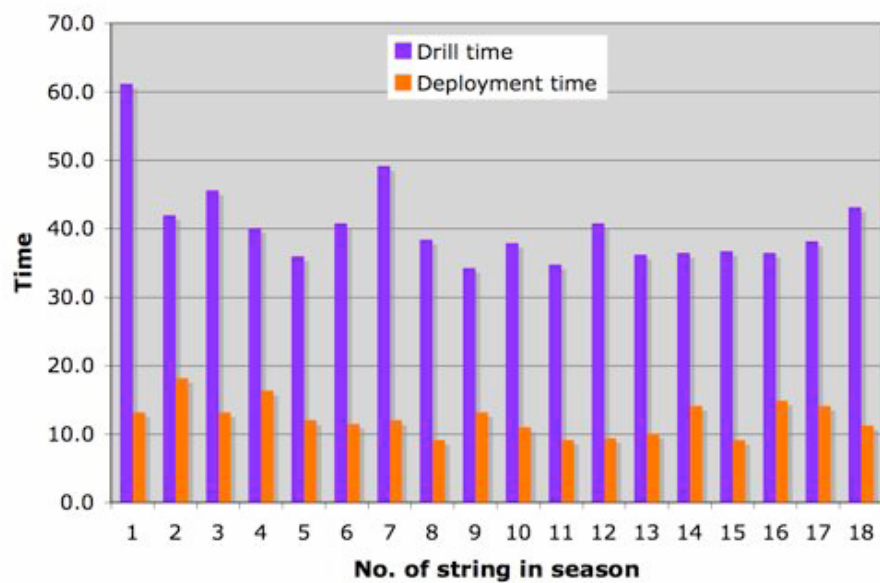


Drilling

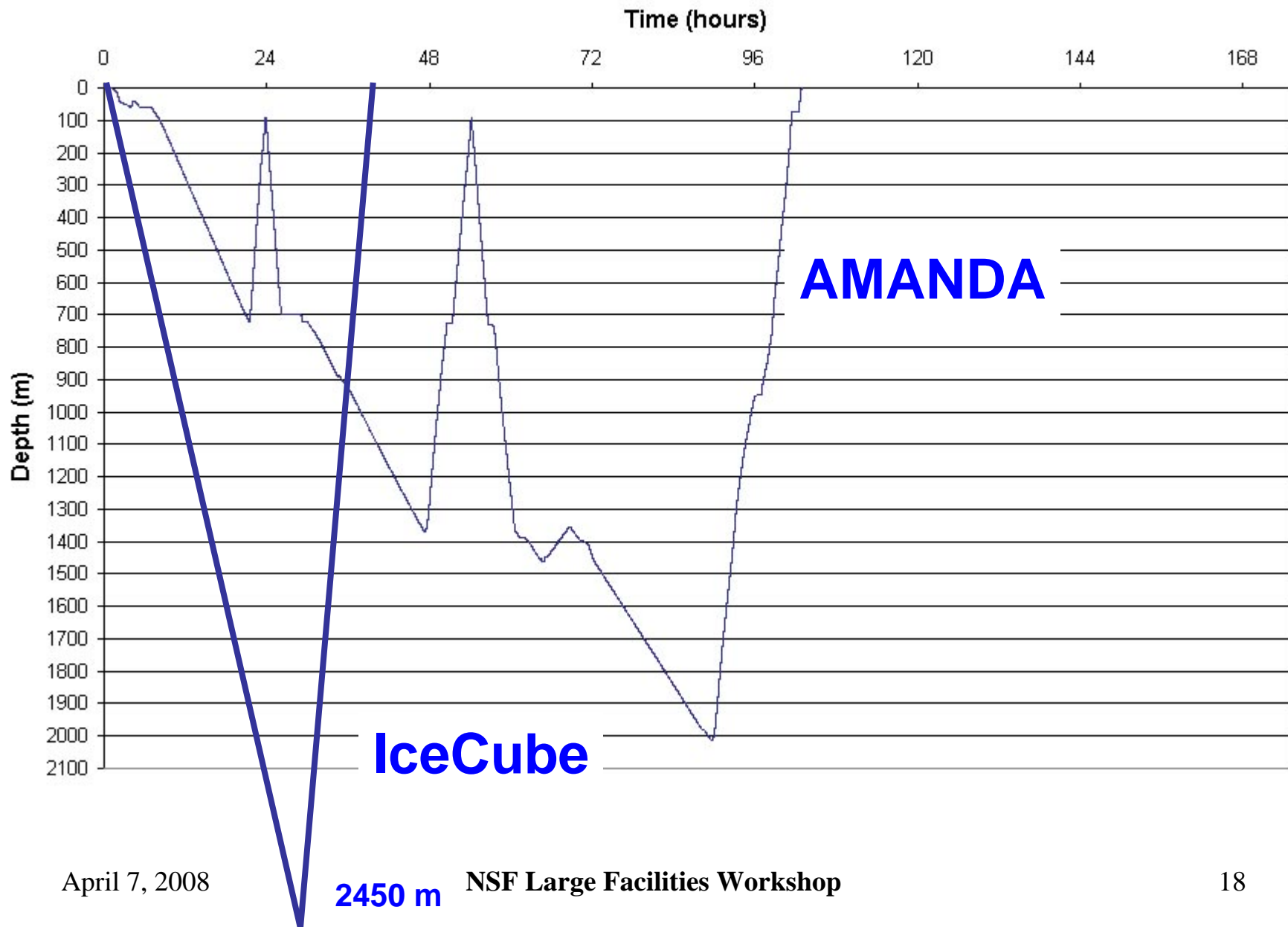
Drilling camp



Drilling and Deployment times 07/08



AMANDA vs. IceCube Drilling



April 7, 2008

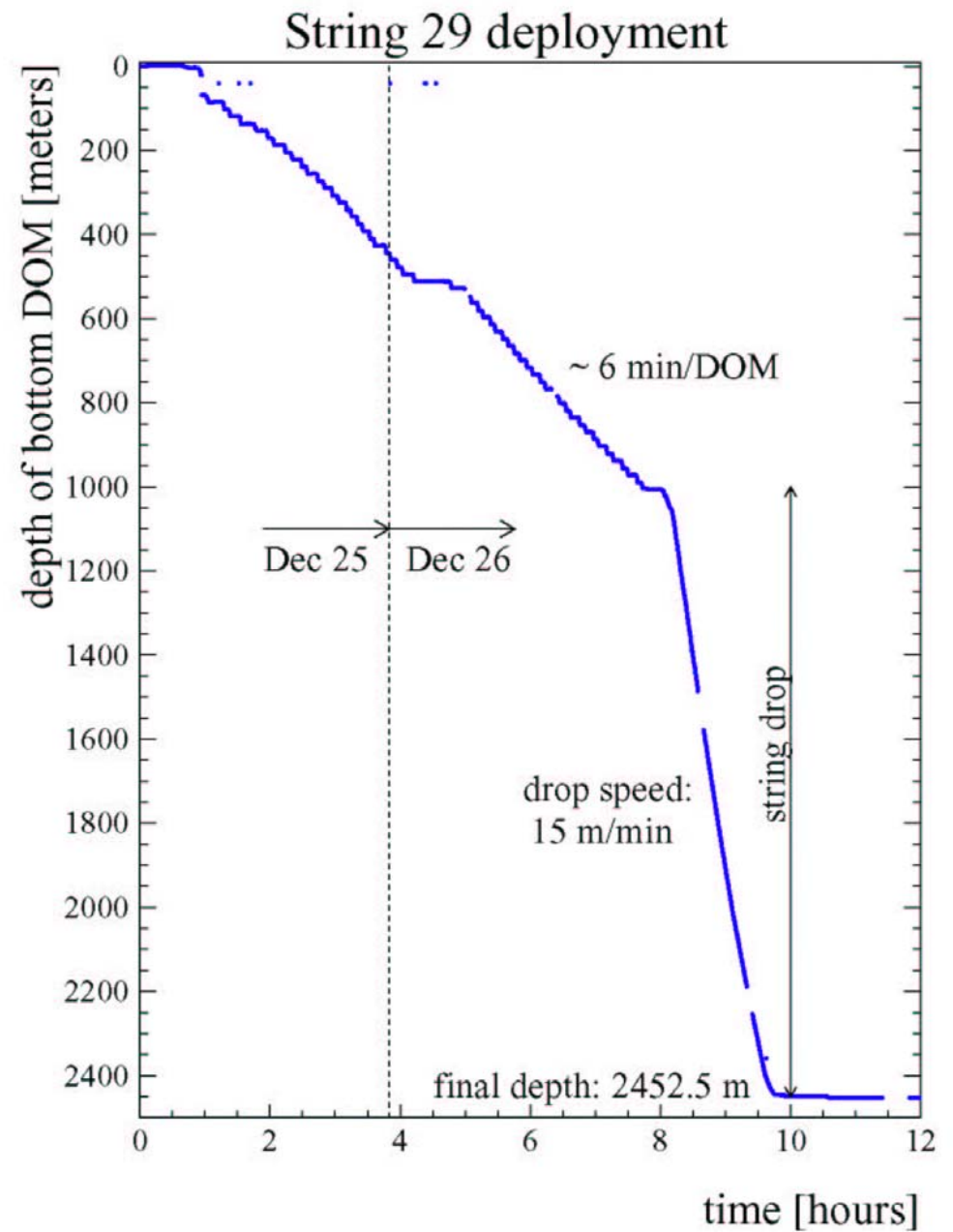
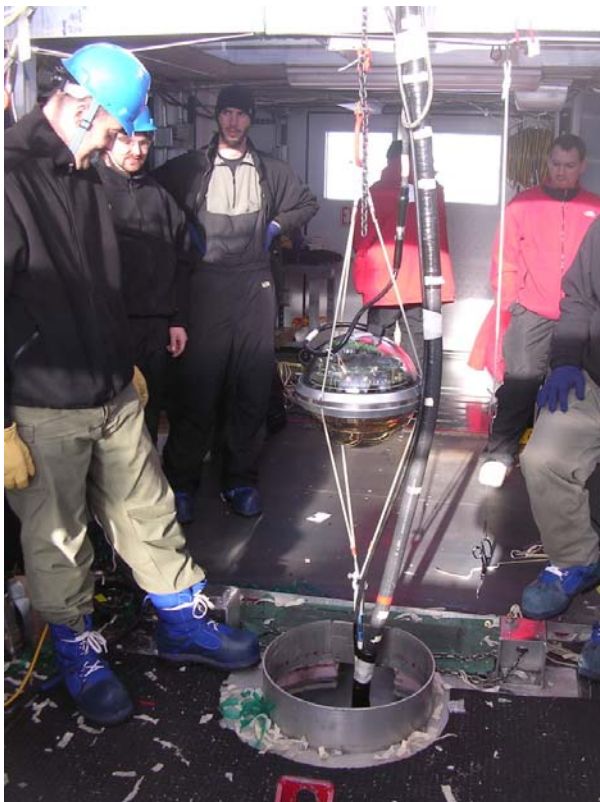
NSF Large Facilities Workshop

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String cable 2500 m

Weight ~6 tons

QuickTime™ and a
Photo - JPEG decompressor
are needed to see this picture.





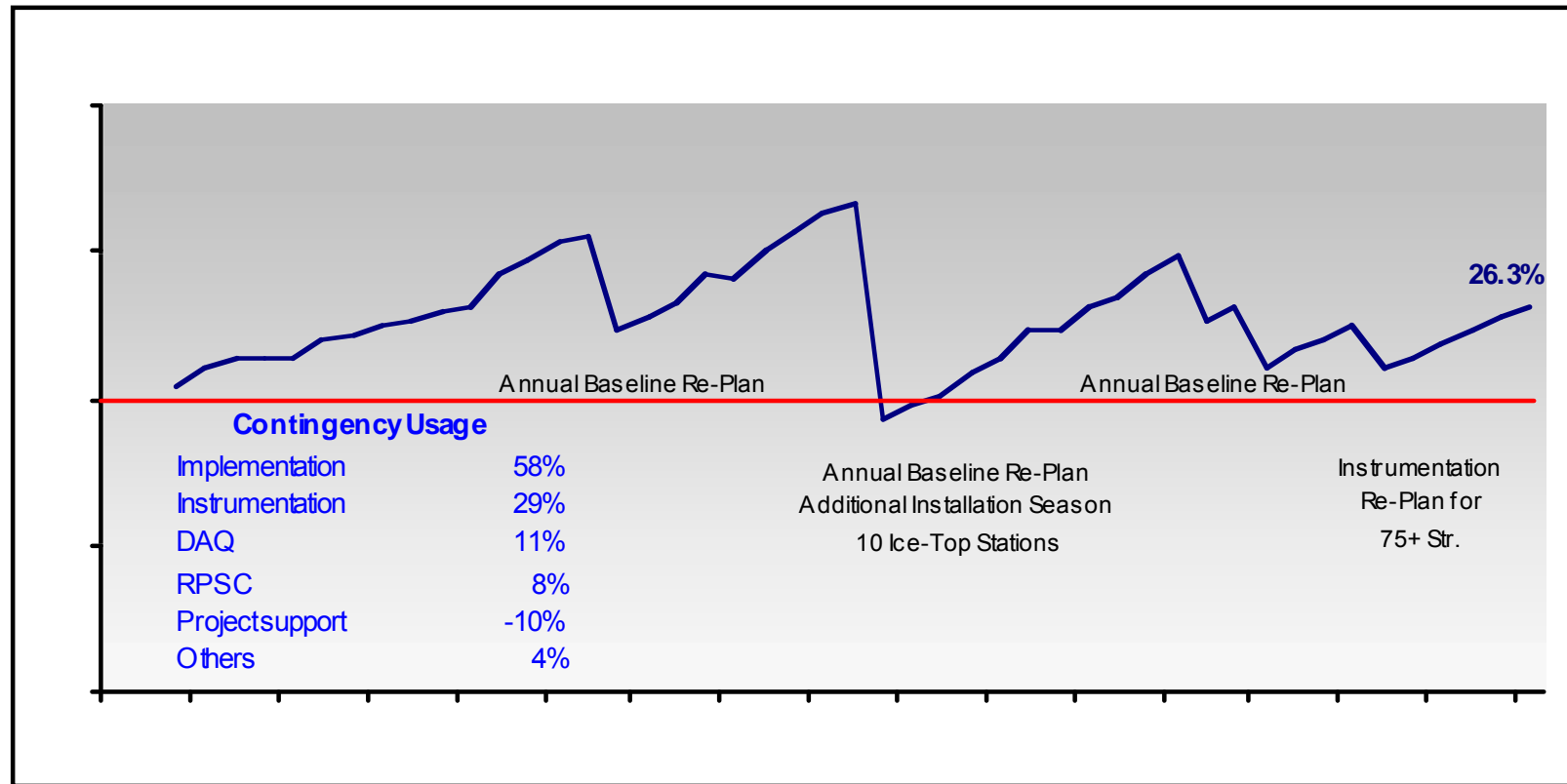


IceCube Laboratory

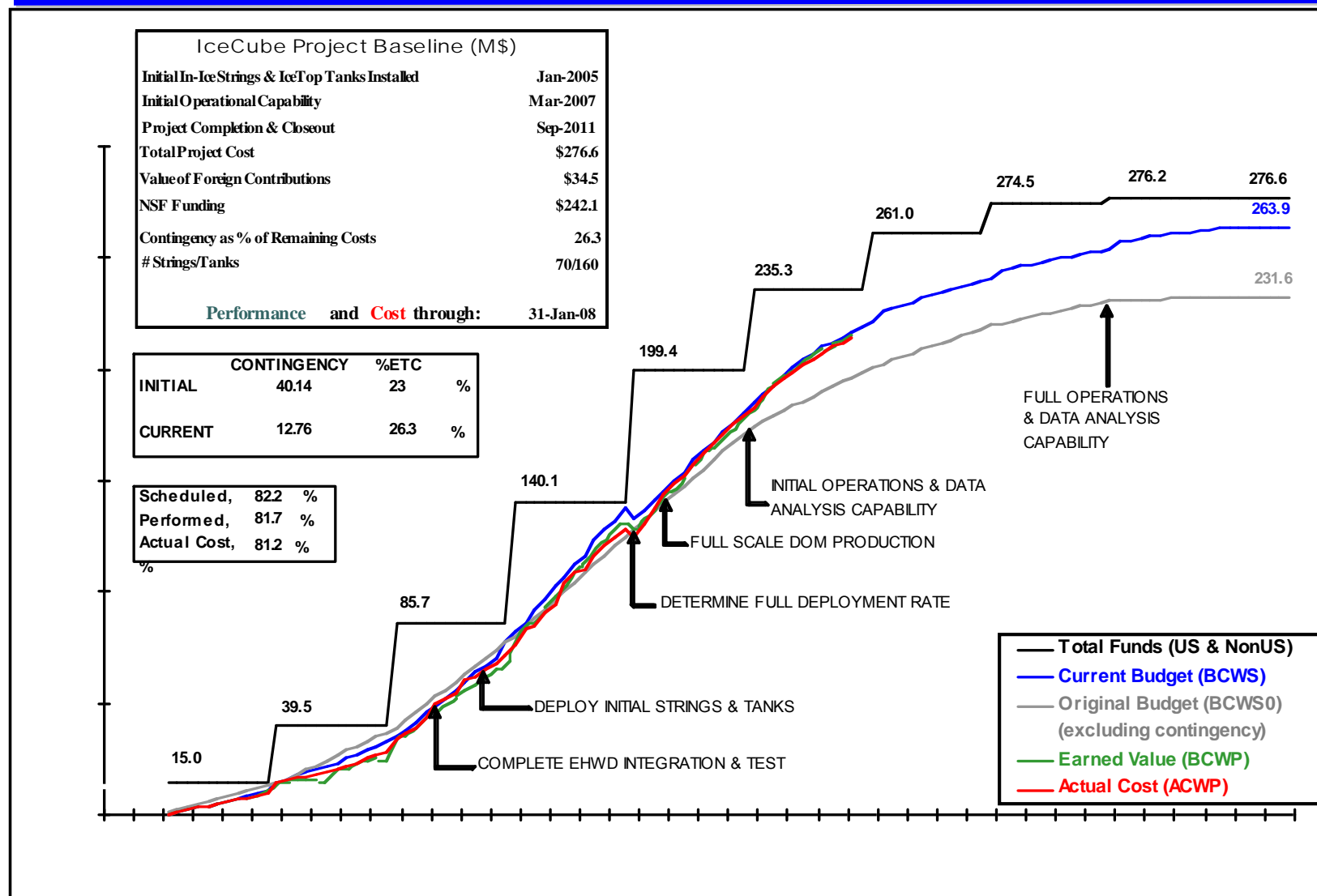


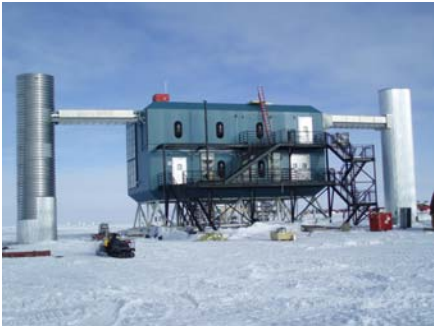
17 racks of computers
Power: 60 kW total for full IceCube
Filtered data sent by satellite

Contingency (% of Work-to-Go)



Cost & Schedule Performance

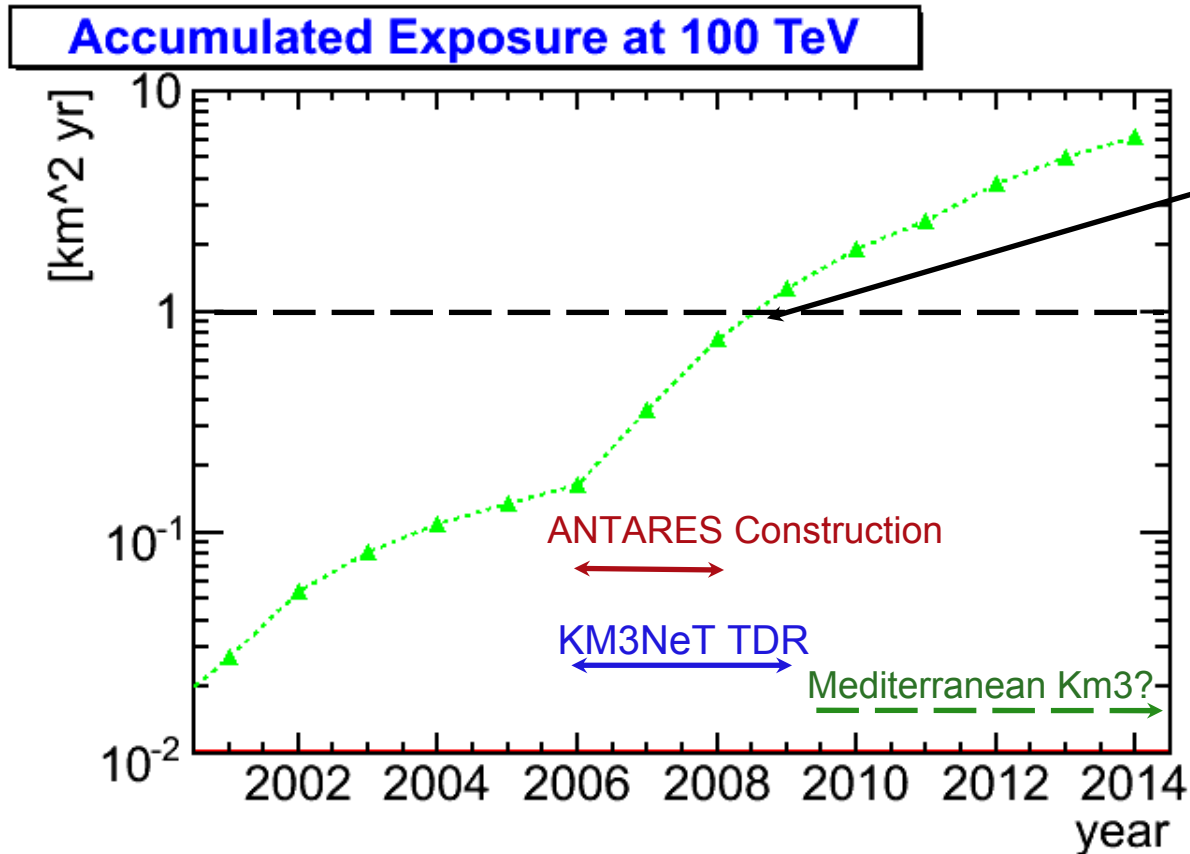




Summary

- Construction project is ~82% complete (earned value)
 - Confident in the quality of completion plans
 - Conservative baseline plan for additional drilling years
 - Materials and labor costs are well understood
 - Continuous attention to safety performance
- Operations formally underway
- Restoring strings as originally planned
 - Plans in place for 80 Strings/IceTop Stations
 - Plan to replace AMANDA with 6+ Strings

Accumulated Exposure



IC40: Science Run starts
in Apr 2008

Effective area for muons at 100 TeV

April 7, 2008

NSF Large Facilities Workshop

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IceCube Challenges

- **NSF's Limited Infrastructure and Experience w/ Large Facilities**
 - Limited experience in “stewardship” role
 - Internal processes developed to meet different needs, typically small grants
 - Operations and transition needs not developed in advance (start-up/ops)
- **Public University Infrastructure and Management**
 - Business and human resource systems designed to meet different needs
 - Schedule imperative not a traditional motivator
- **Detector Size and Collaboration Complexity**
 - Large extrapolation from AMANDA (prototype) to IceCube detector
 - Increase in number of institutions and dependencies



IceCube Opportunities

- USAP Infrastructure and Support
 - Established site and supply chain (NSF, Raytheon, USAF, NYANG)
- University of Wisconsin Commitment and Infrastructure
 - Management support and overhead resources
 - Physical Sciences Laboratory (PSL) staff and facilities
 - Willingness to adapt university systems to meet large project needs
- Instrumentation production in 3 locals - U.S., Germany, Sweden
- Annual Feedback Loop - seasonal goals vs. accomplishments
- Early Operations and Research
 - Opportunity to address detector performance and long term support issues



IceCube Experience

- Construction cost baseline set realistic goals
 - Original scope reduced from 80 to 70 strings plus 20% contingency
 - Included plan for early start of operations
- NSF
 - NSF collaborated with UW to find ways to best support the project
 - Dedicated IceCube Project Officer
 - Annual Peer Reviews and Action Tracking
 - Construction funding secure and predictable
 - Operations and research funding more uncertain
- UW Resources and People
 - University leadership involvement critical at key junctures
 - Experienced personnel essential to project success



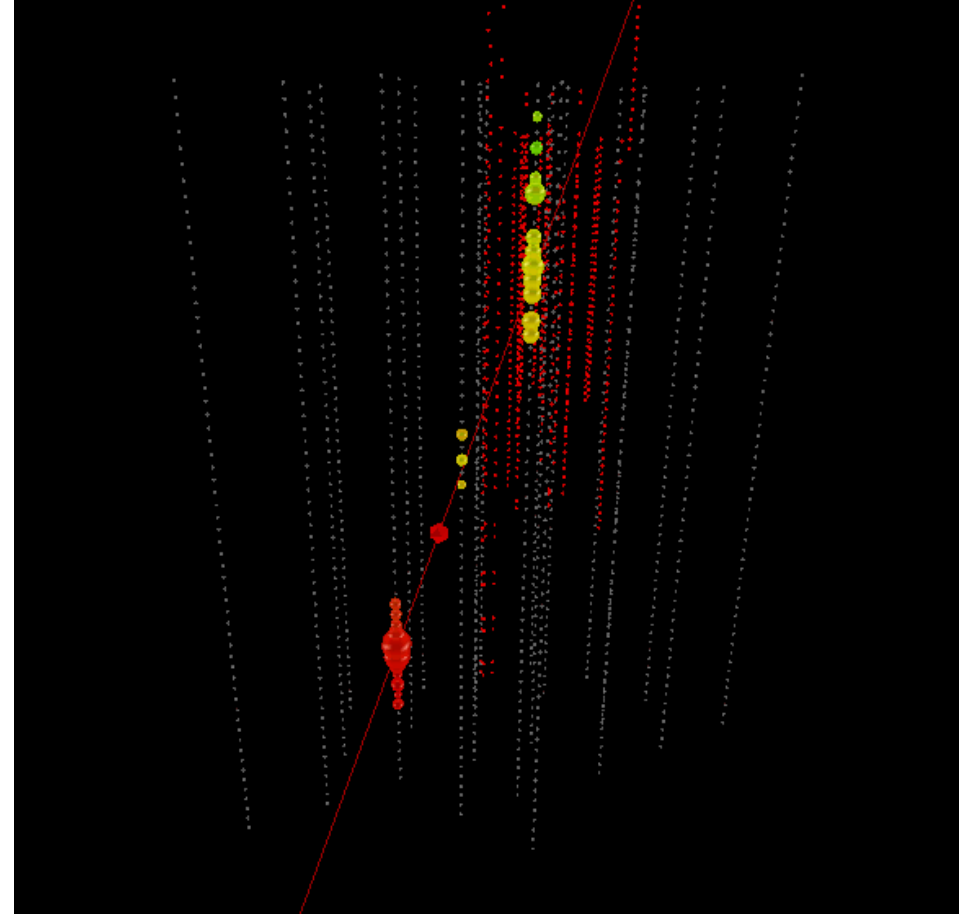
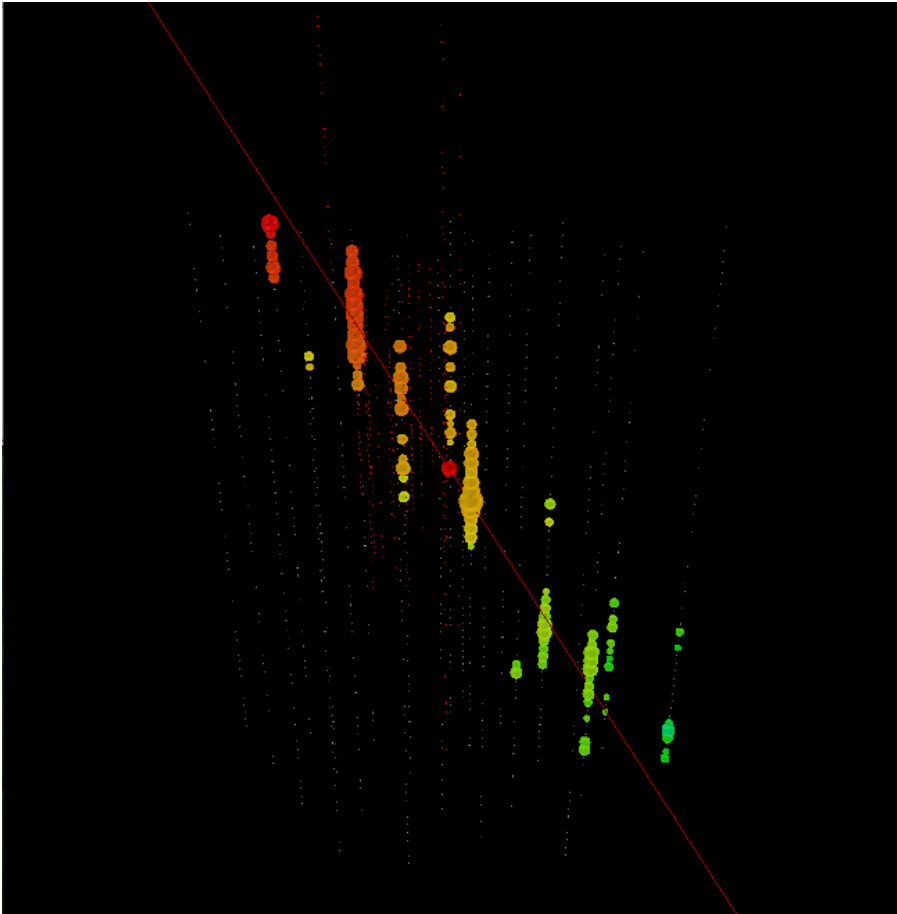
General Observations

- Ingredients to Success
 - Strong host role
 - Populate the organization with high quality people
 - Project & Collaboration Leadership
 - Makes decisions - seeking consensus whenever possible
 - Serves as an umbrella to the project team
 - Manages expectations and communicates plans and results
 - Develops strategies and revises goals as needed
 - Establish realistic project goals
 - Maintain credibility with stakeholders
 - Seek collective ownership of problems and solutions

IC22 Events

(Red hits = early; yellow/green/blue = later)

IceCube DOM locations blue, AMANDA OM locations red



Downward cosmic-ray event (“muon bundle”)

April 7, 2008

Upward candidate ν event

NSF Large Facilities Workshop

IC40 Event

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